

Modelling the Milky Way exoplanet population

Chloé Padois, Daniel del Ser, Friedrich Anders, et al.



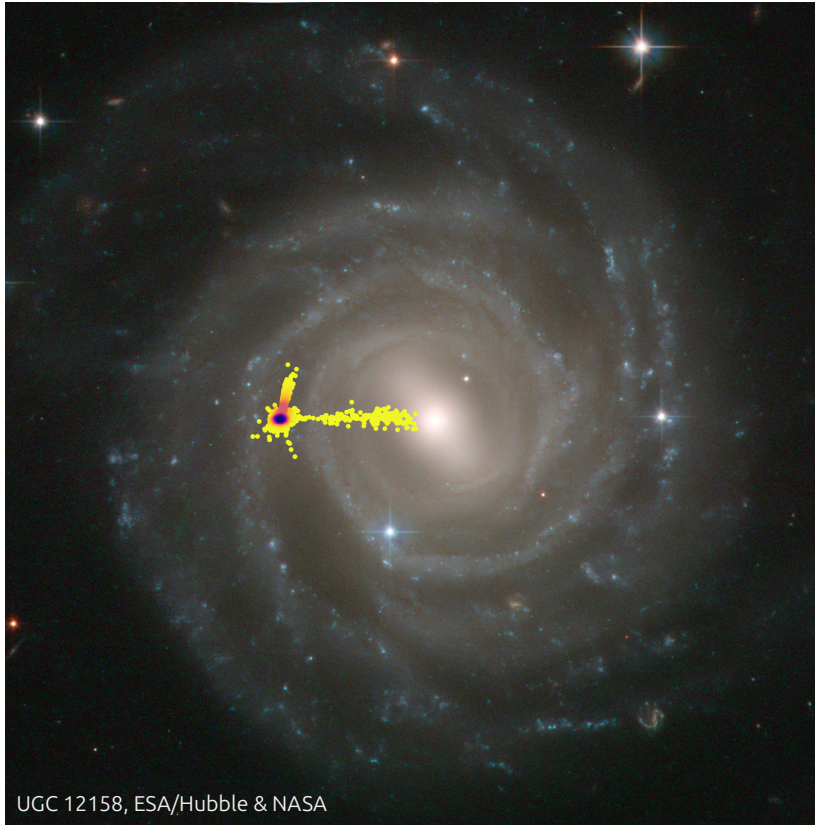
MWGaiaDN

This project is a Horizon Europe Marie Skłodowska-Curie Actions Doctoral Network funded under grant agreement no. 101072454.



Context: study exoplanets from a galactic point of view

- Known exoplanets: observational bias + **restricted to the SN**

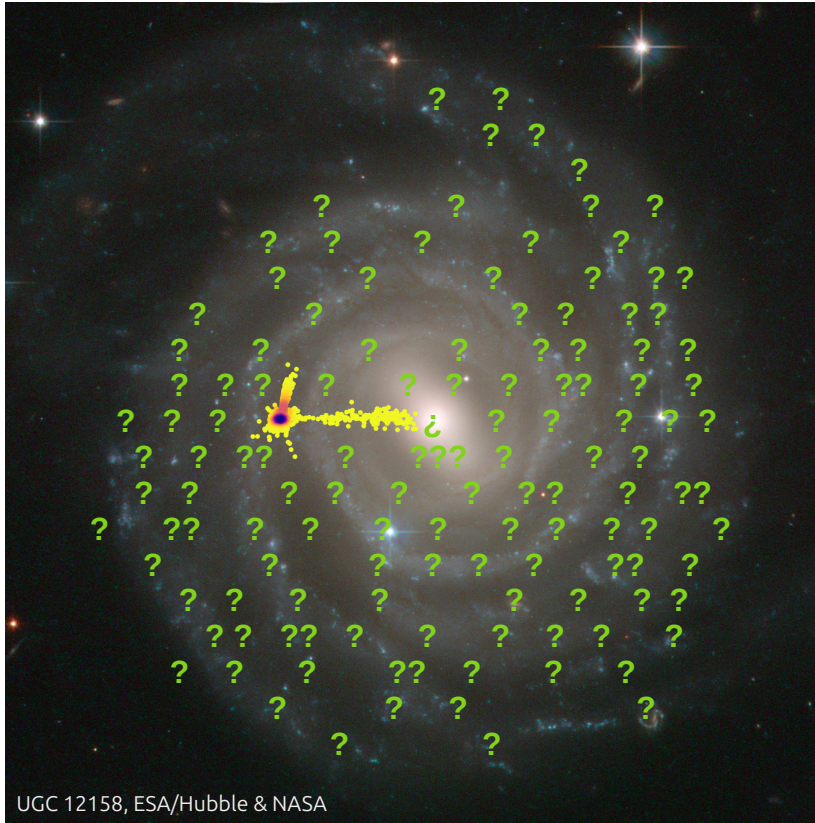


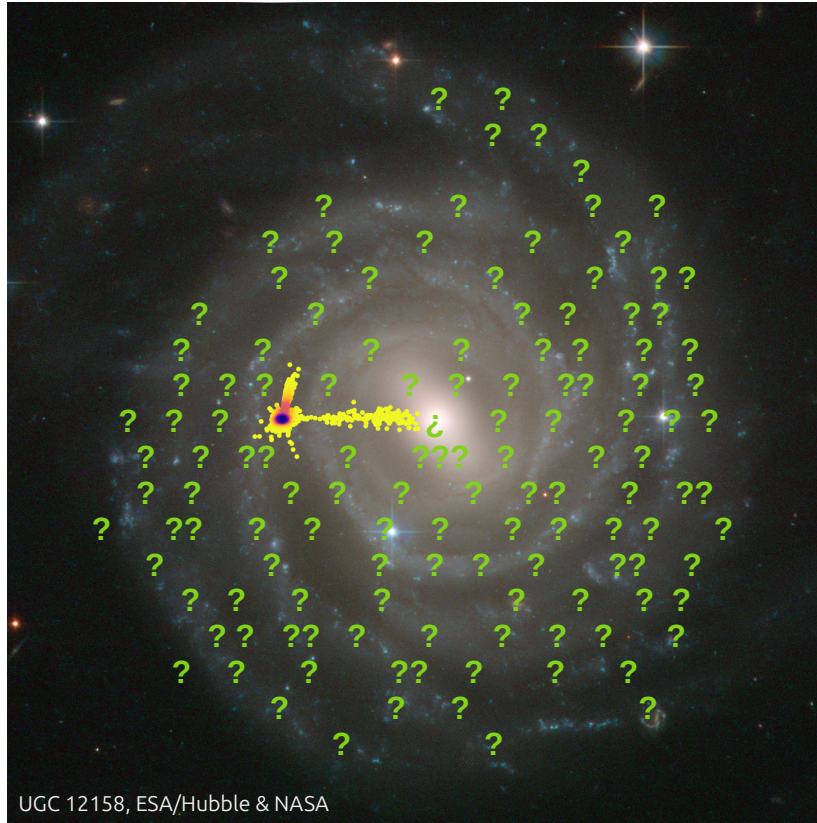
Context: study exoplanets from a galactic point of view

- Known exoplanets: observational bias + restricted to the SN

Goal: simulate the (unbiased) exoplanet population in the MW

- Exoplanets everywhere? Homogeneously distributed?
- Changes with galactic evolution?





Context: study exoplanets from a galactic point of view

- Known exoplanets: observational bias + restricted to the SN

Goal: simulate the (unbiased) exoplanet population in the MW

- Exoplanets everywhere? Homogeneously distributed?
- Changes with galactic evolution?

Method: Framework to quickly generate exoplanet population

- Around any stellar population (mass and metallicity)
- Combining observations and planetary formation models
- Predict yields of future missions → **constrain models**

1) General methodology

2) Comparing with observations

Generate exoplanet population



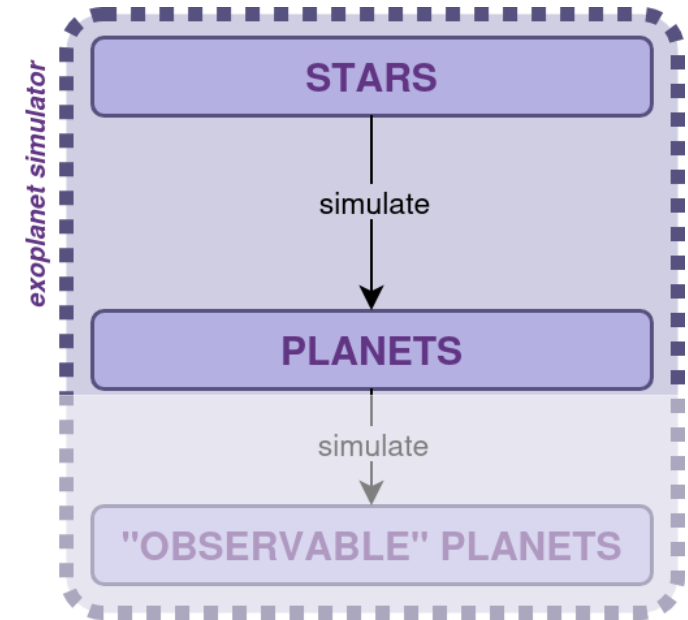
Evaluate our results:

Detectability by *Kepler*

Predictions for future surveys

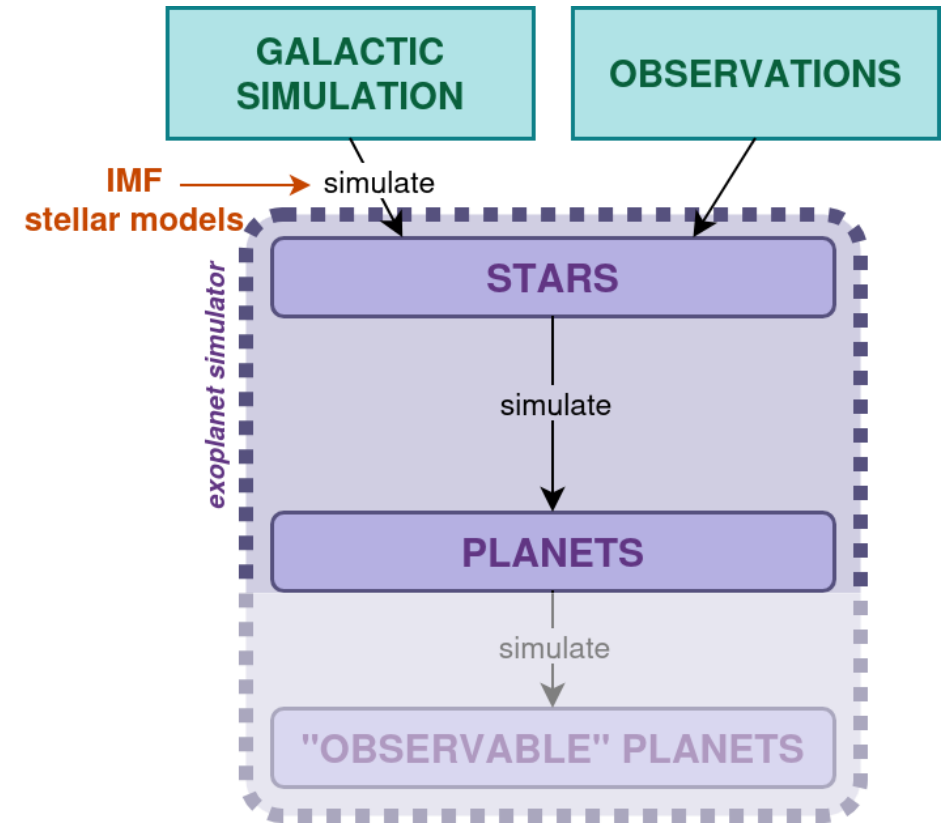
1) General methodology

Generate exoplanet population



1) General methodology

Generate exoplanet population

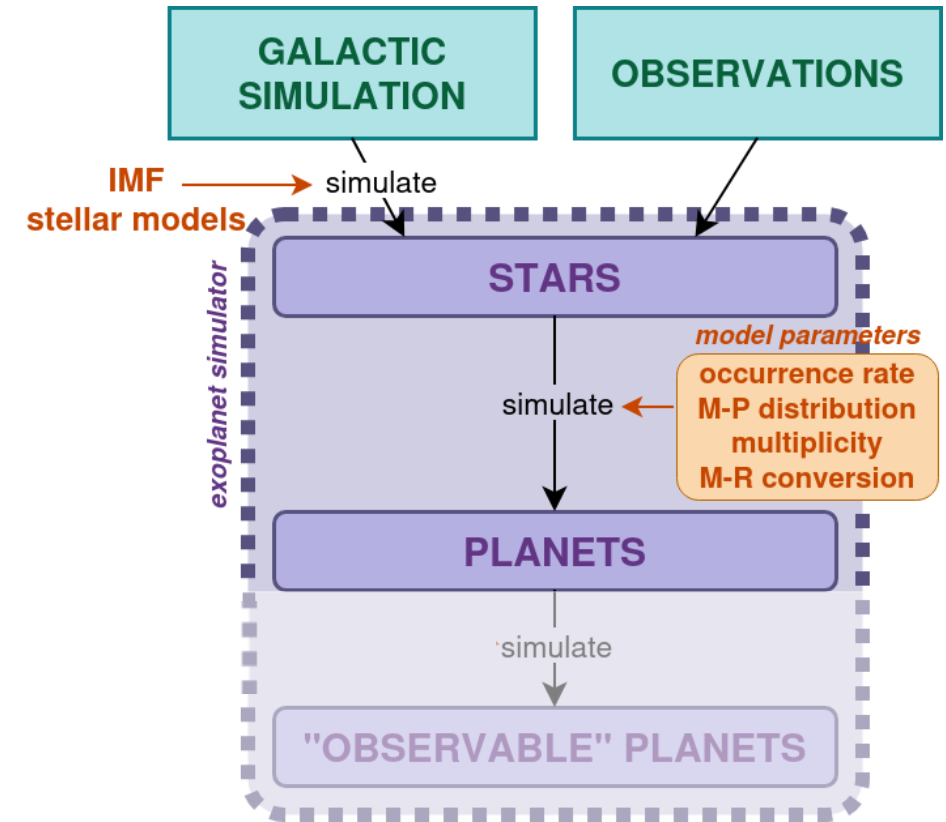


1) General methodology

Generate exoplanet population

Occurrence rates and multiplicity

Assign physical parameters



General methodology: Creation of a synthetic exoplanet population

Occurrence rates and multiplicity

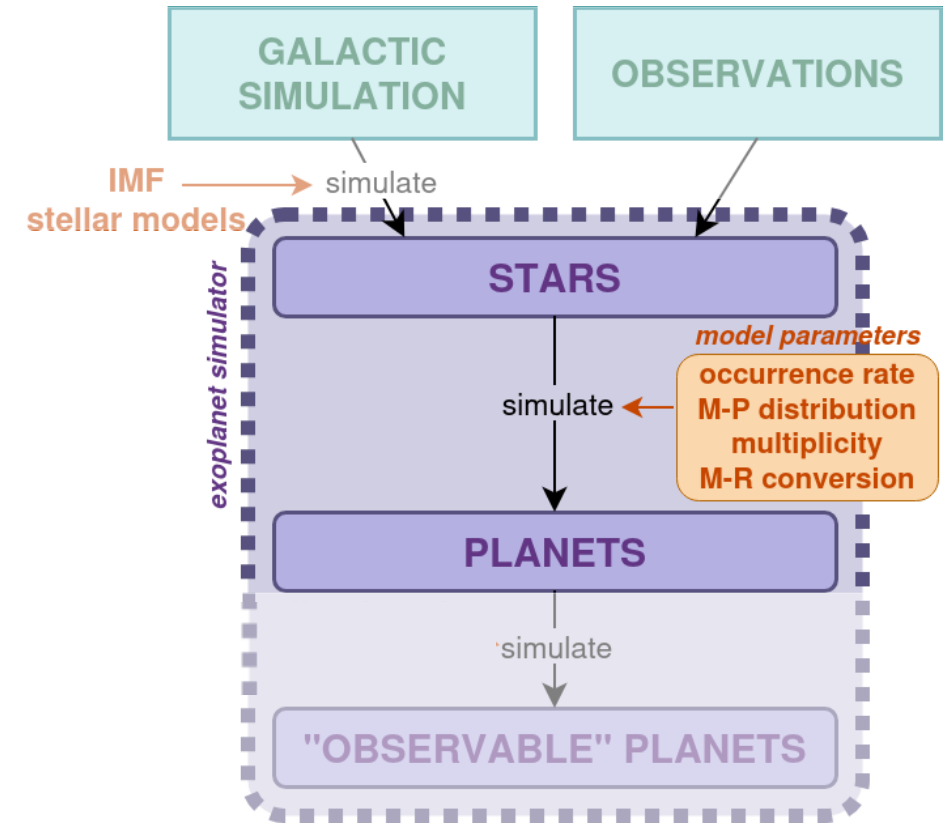
- Probability for a star to harbour planets

→ $f(M^*, [Fe/H]^*)$

Earth-like

Super-Earth + Neptunians

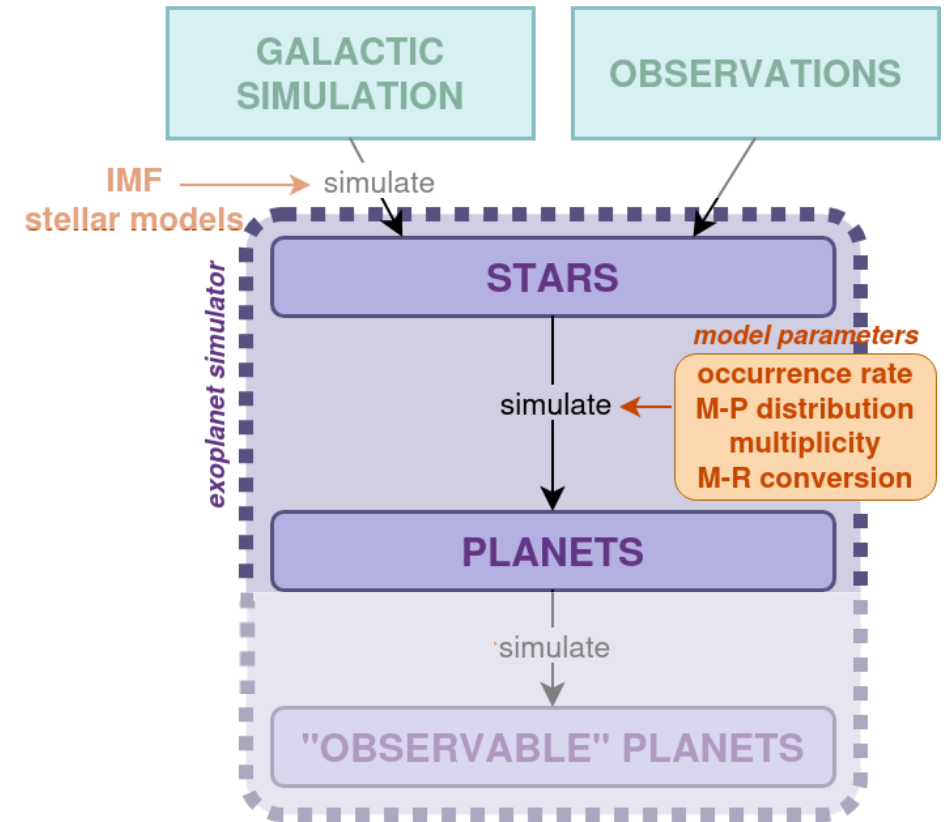
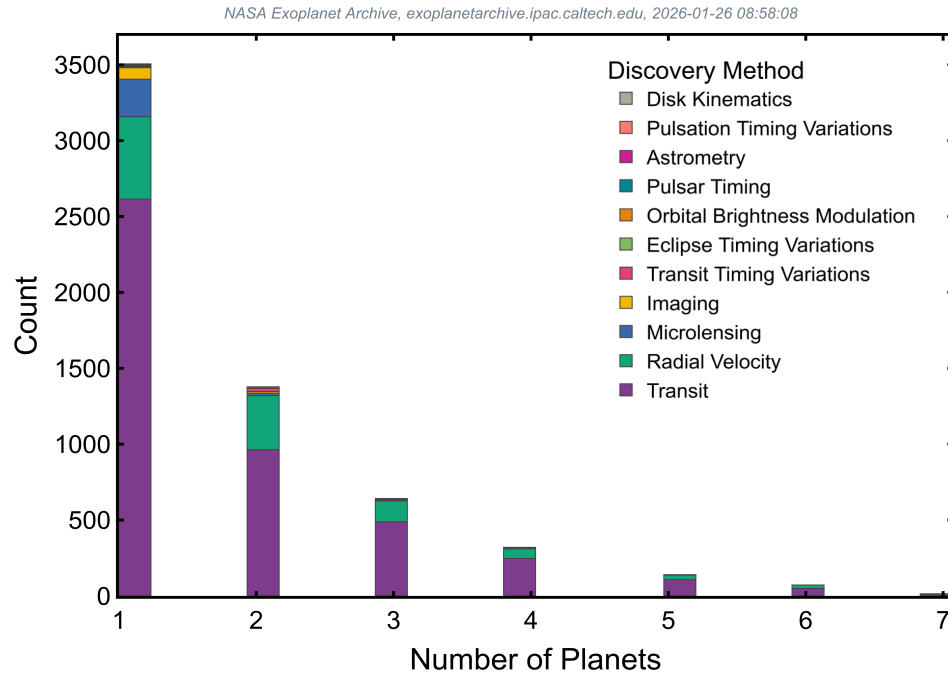
Sub-Giants + Giants



General methodology: Creation of a synthetic exoplanet population

Occurrence rates and multiplicity

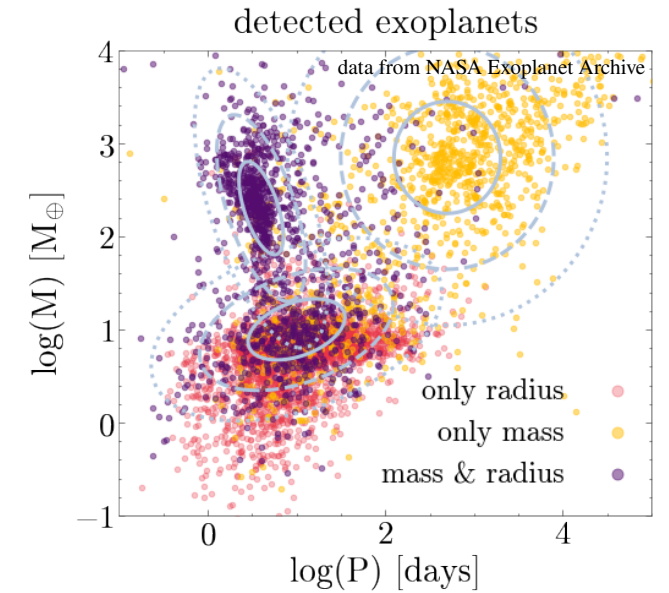
- Probability for a star to harbour planets
 → $f(M^*, [Fe/H]^*)$
- How many? → multiplicity: $f(M^*)$



General methodology: Creation of a synthetic exoplanet population

Assign physical parameters

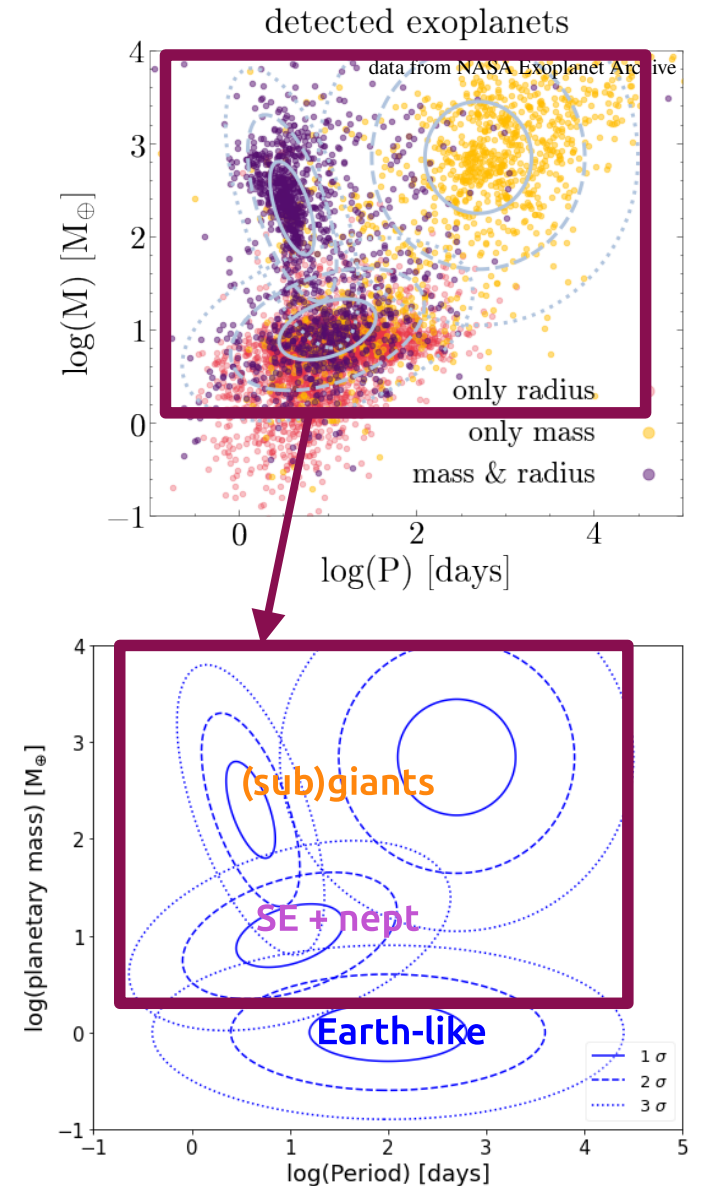
- Distribution in mass-period diagram



General methodology: Creation of a synthetic exoplanet population

Assign physical parameters

- Distribution in mass-period diagram:
Combining **observations** and **planetary formation model**
(planetesimal accretion synthesis, from Drążkowska+23)

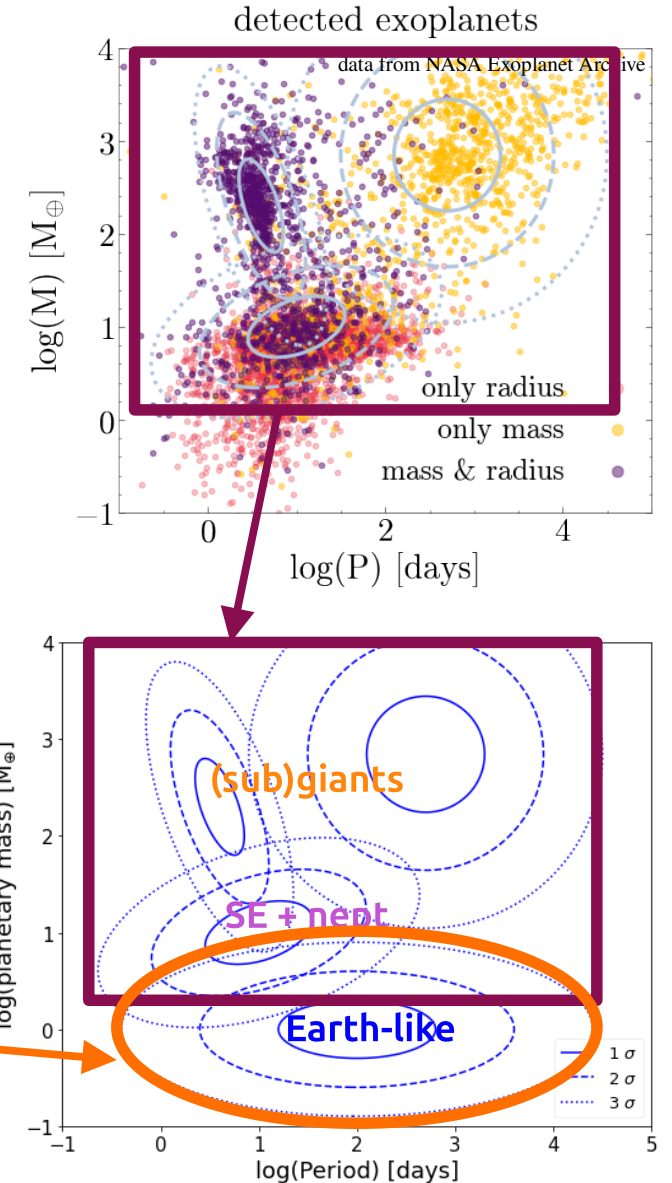
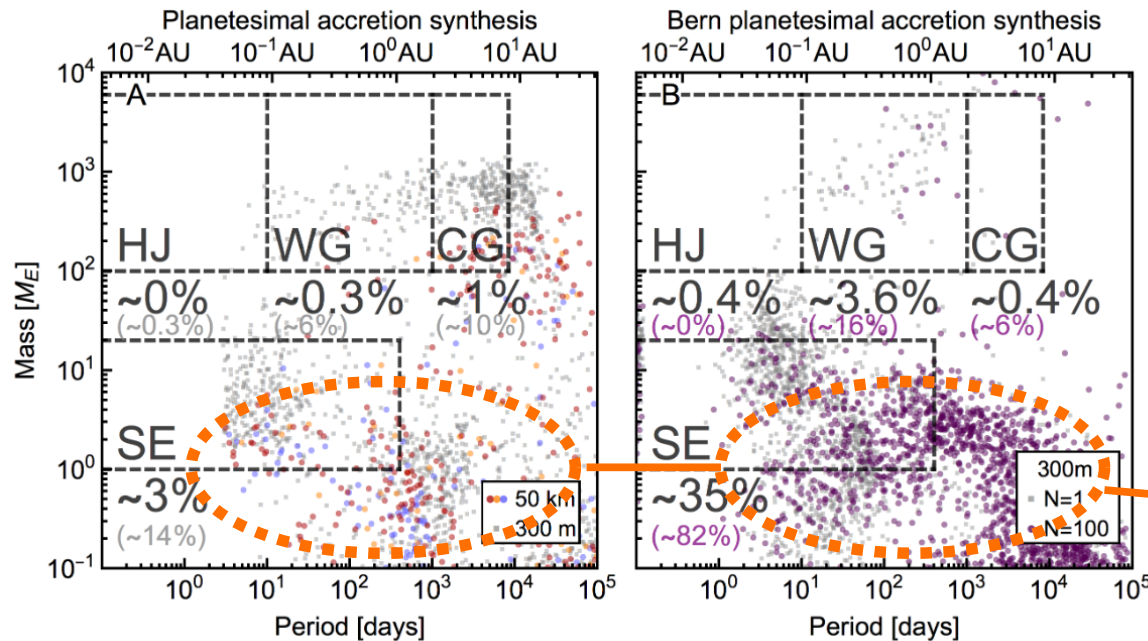


Modelling the Milky Way exoplanet population

General methodology: Creation of a synthetic exoplanet population

Assign physical parameters

- Distribution in mass-period diagram: Combining **observations** and **planetary formation model** (planetesimal accretion synthesis, from Drążkowska+23)

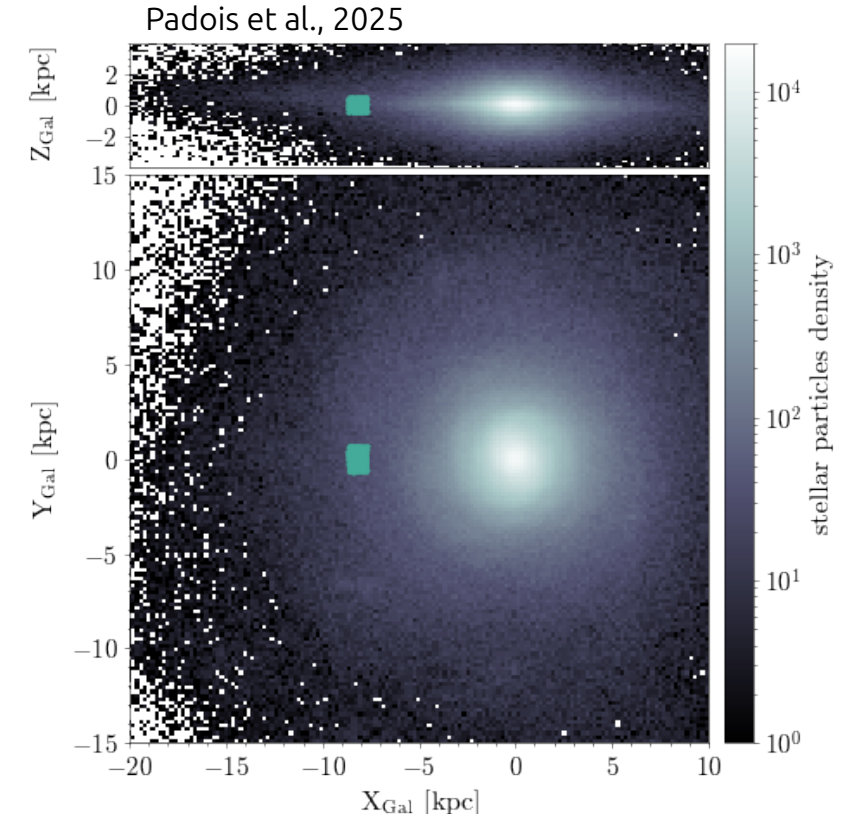


Modelling the Milky Way exoplanet population

General methodology: Creation of a synthetic exoplanet population

Some results: the Solar Neighbourhood

8.3M stars > 3.4M singles → 2.9 million planetary systems
→ **22.5 million planets**

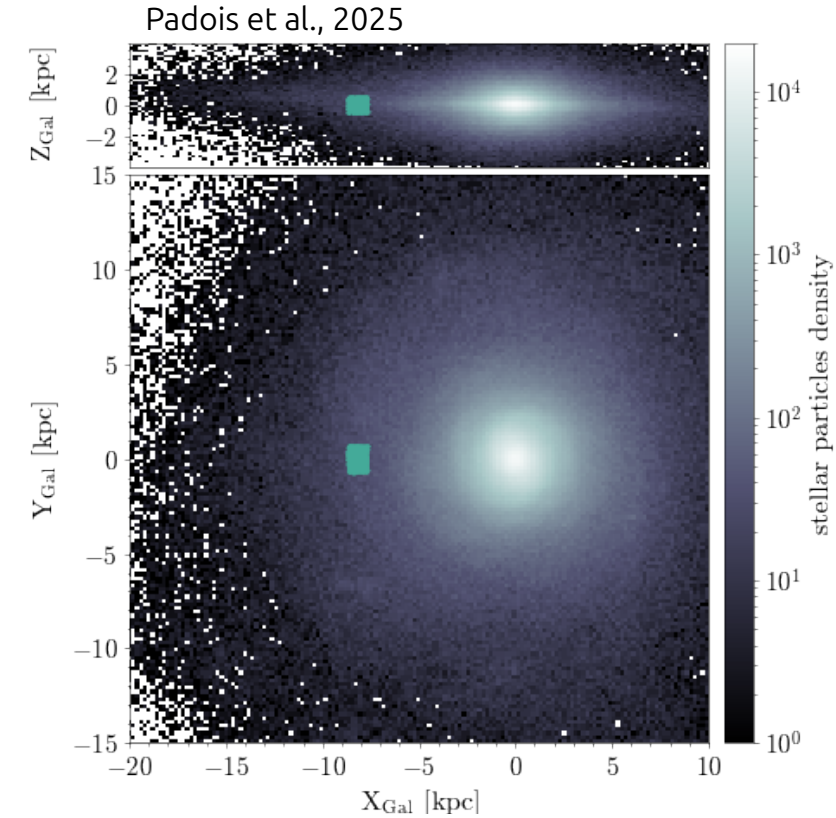
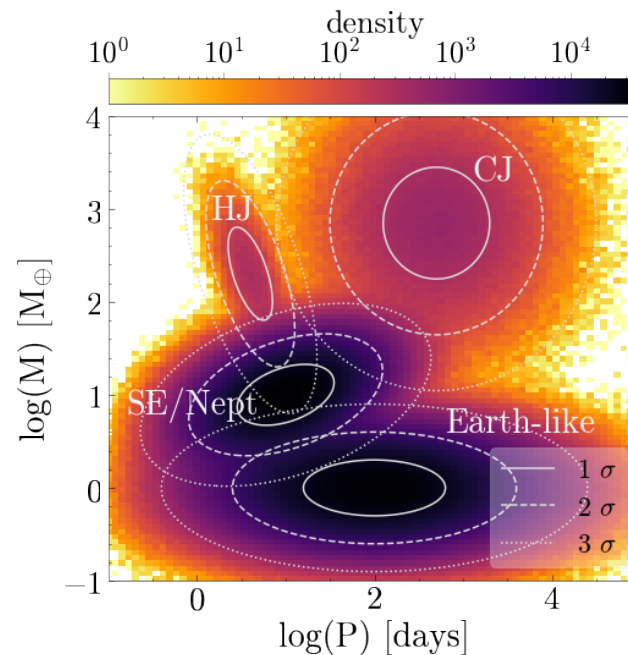
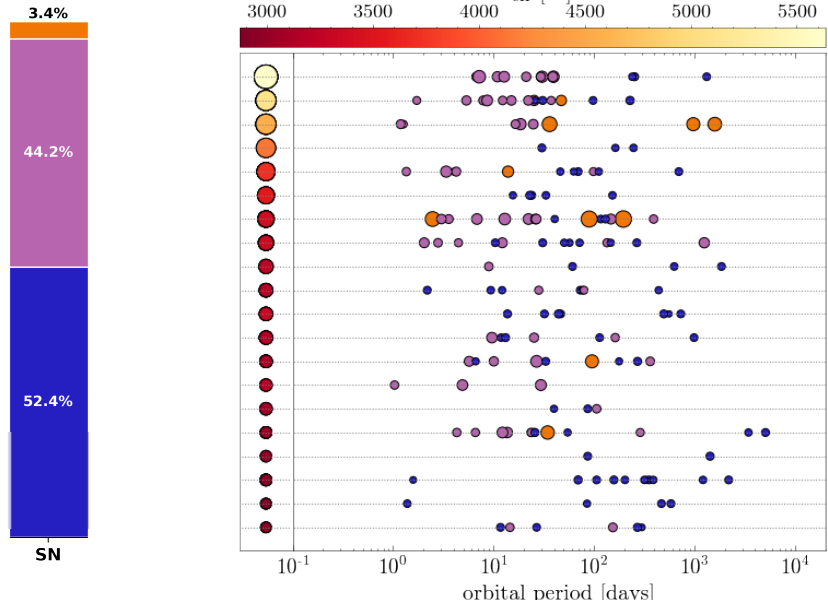


Modelling the Milky Way exoplanet population

General methodology: Creation of a synthetic exoplanet population

Some results: the Solar Neighbourhood

8.3M stars > 3.4M singles → 2.9 million planetary systems
→ **22.5 million planets**



Earth-like

Super-Earth + Neptunians

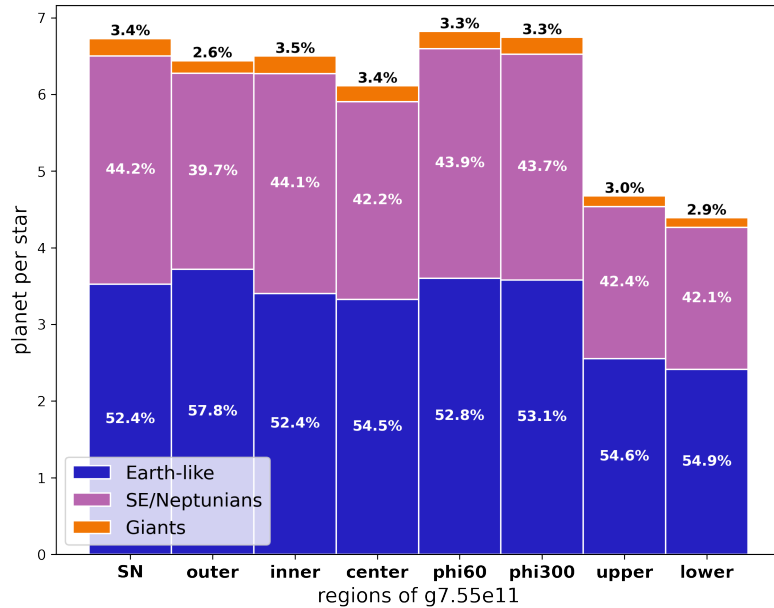
Sub-Giants + Giants

Modelling the Milky Way exoplanet population

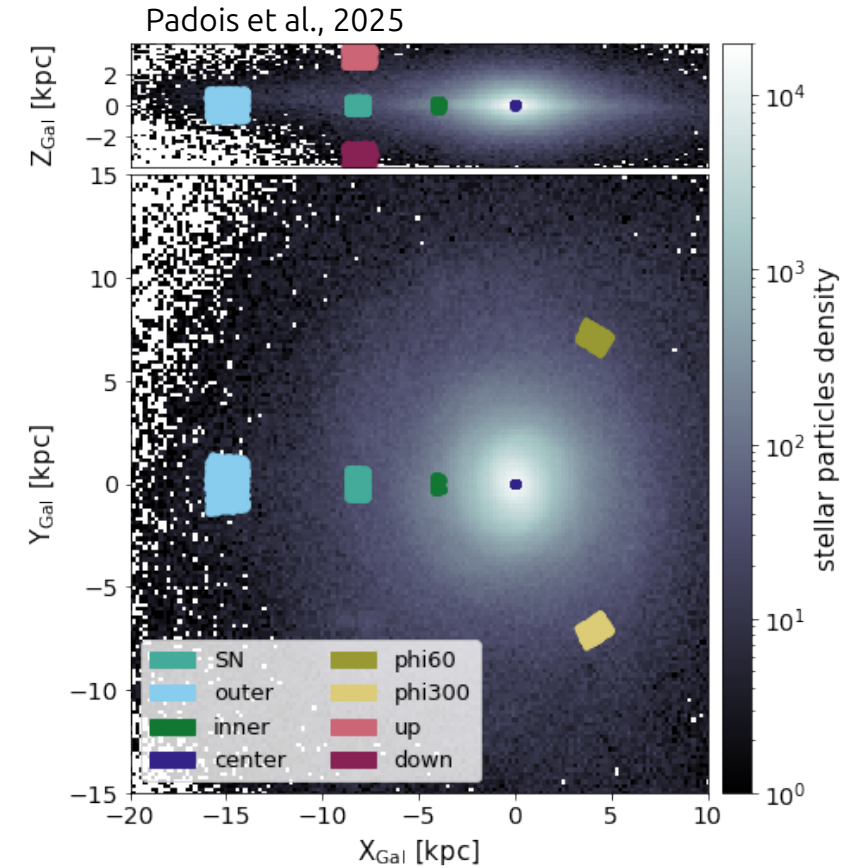
General methodology: Creation of a synthetic exoplanet population

Some results: galactic environment

→ halo regions host less exoplanet per star



→ strongly model-dependent
→ How realistic is this simulated exoplanet population?



1) General methodology

2) Comparing with observations

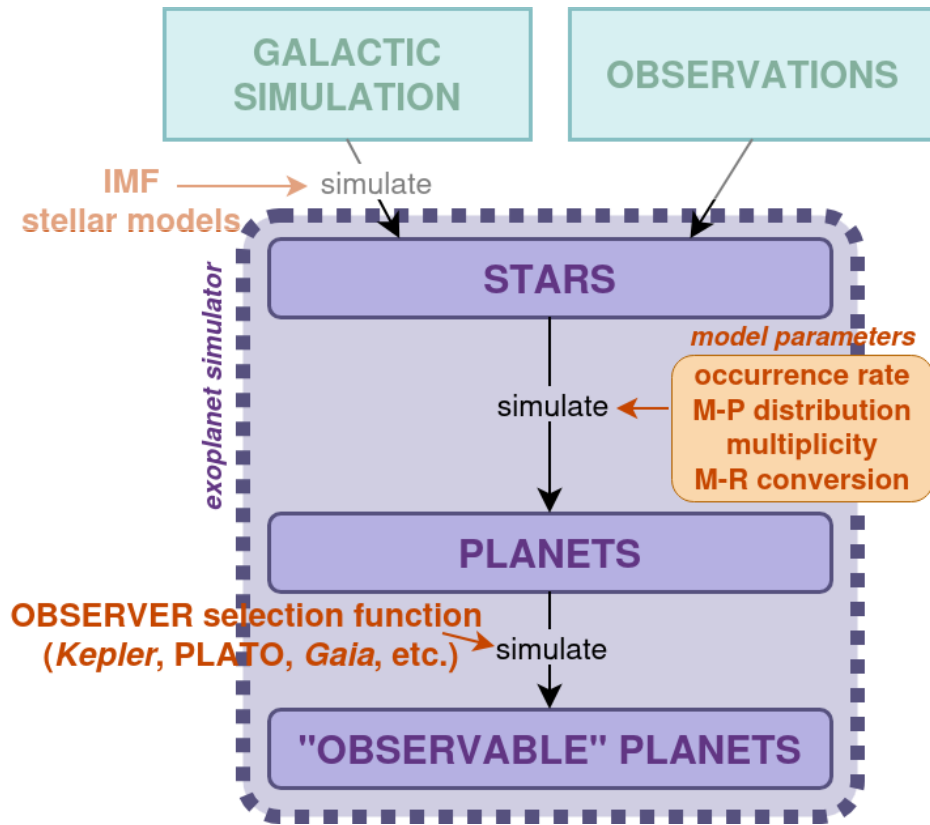
Generate exoplanet population



Evaluate our results:

Detectability by *Kepler*

Predictions for future surveys



2) Comparing with observations

Evaluate our results:

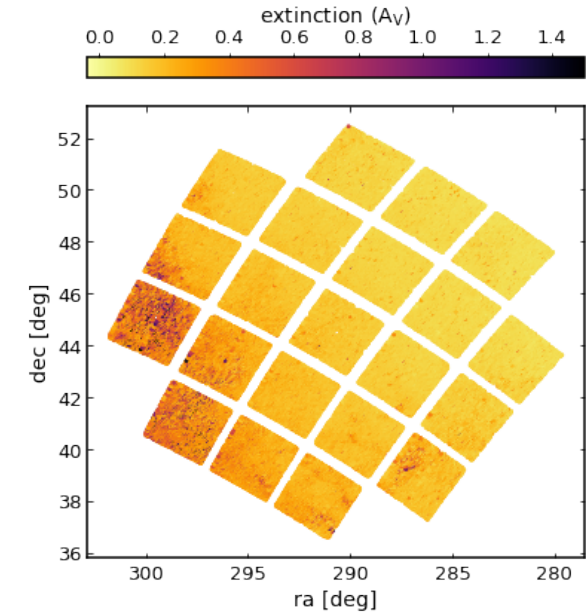
Detectability by *Kepler*

Predictions for future surveys

Comparing with observations: evaluate our results

Kepler field of view

- Apply same process → *Kepler* target list (from Wolniewicz+21)
→ 1.6M planets simulated around 174k stars
- Assign random inclination
- Reproduce *Kepler* observation: compute SNR



Comparing with observations: evaluate our results

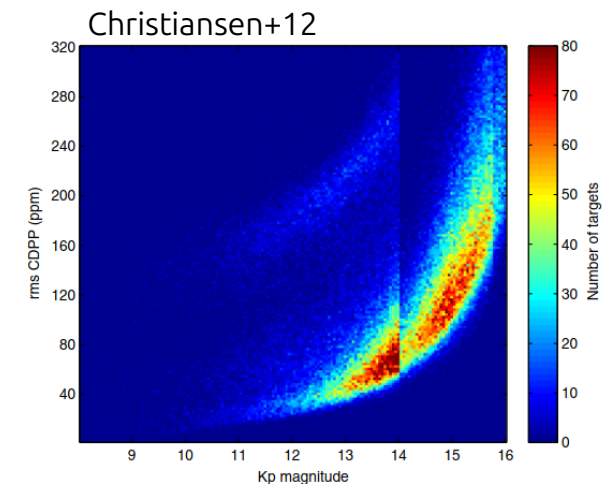
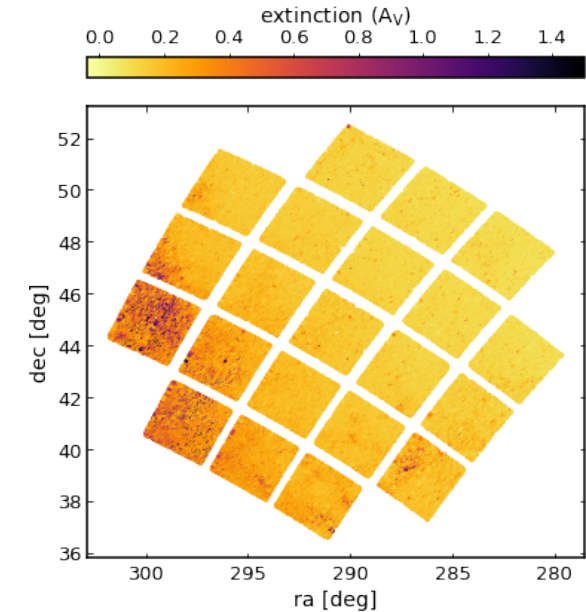
Kepler field of view

- Apply same process → *Kepler* target list (from Wolniewicz+21)
→ 1.6M planets simulated around 174k stars
- Assign random inclination
- Reproduce *Kepler* observation: compute SNR

$$\text{SNR} = \sqrt{N_{\text{tr}}} \times \frac{\delta}{\text{CDPP}_{\text{eff}}}$$

Annotations:

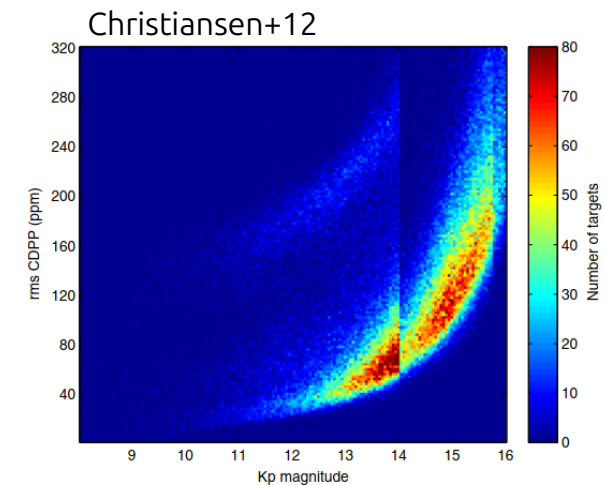
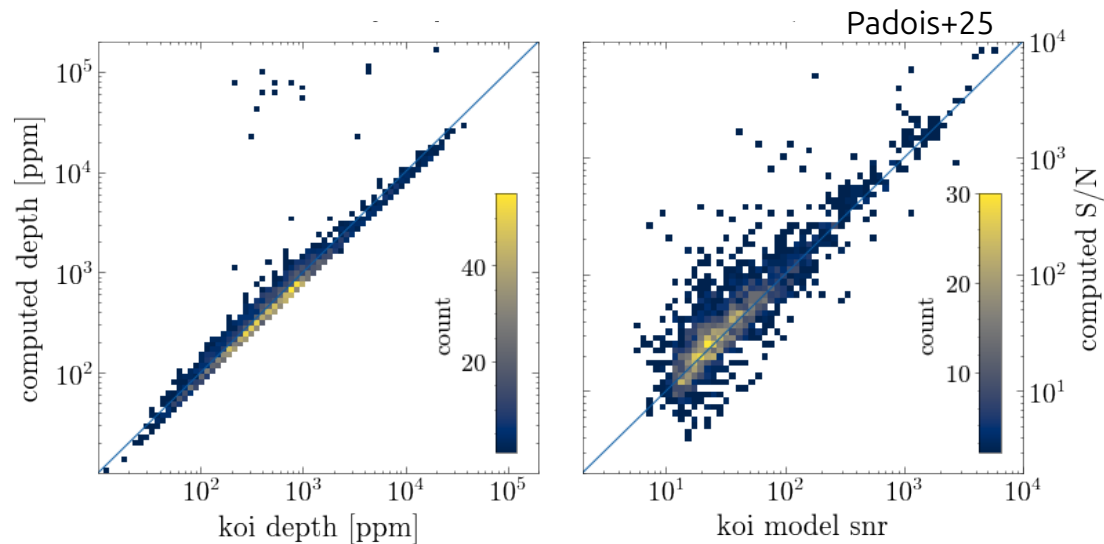
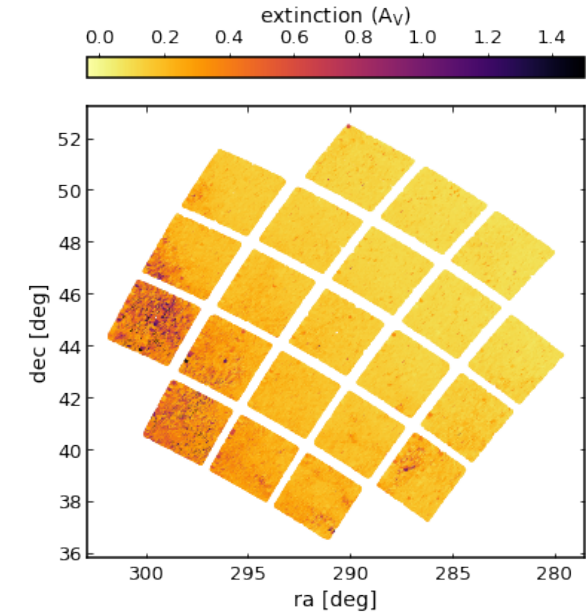
- period (points to $\sqrt{N_{\text{tr}}}$)
- radius (points to δ)
- stellar magnitude (points to CDPP_{eff})



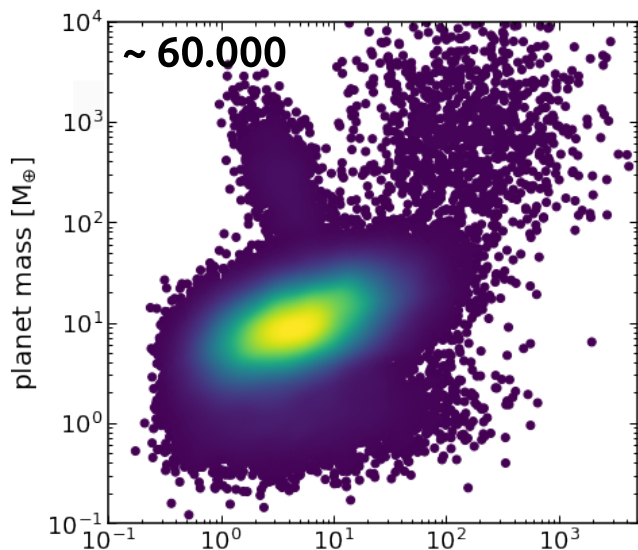
Comparing with observations: evaluate our results

Kepler field of view

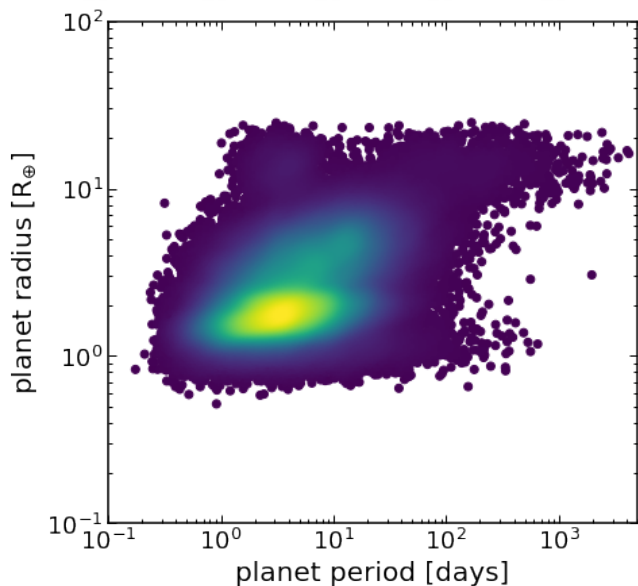
- Apply same process → *Kepler* target list (from Wolniewicz+21)
→ 1.6M planets simulated around 174k stars
- Assign random inclination
- Reproduce *Kepler* observation: compute SNR



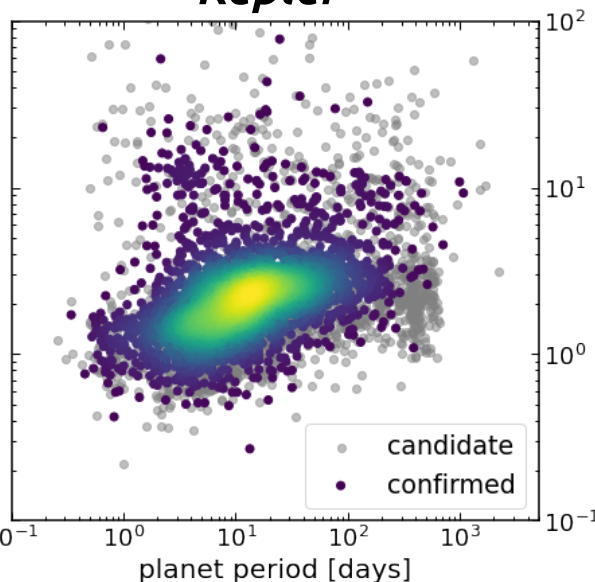
Comparing with observations: evaluate our results



- Conversion from mass to radius (based on Parc+24)
- Over-estimation of the number of “detected” exoplanets
- Differences in the R-P distribution



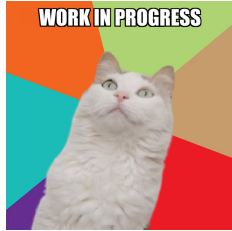
Kepler



→ Adjust model parameters:

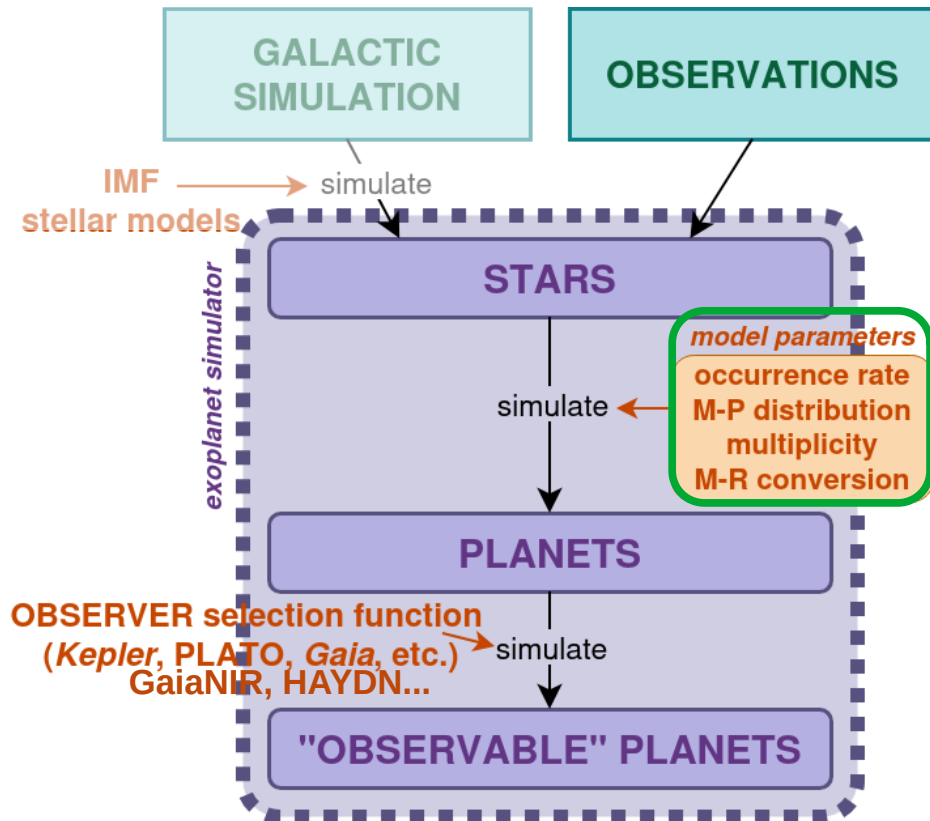
- M-P distributions
- occ. rate, multiplicity, ...

Comparing with observations: adjust model parameters

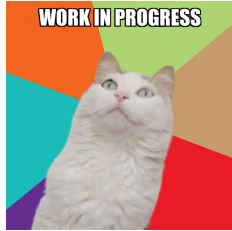


Simulation-Based Inference (SBI)

- Variation of the model parameters

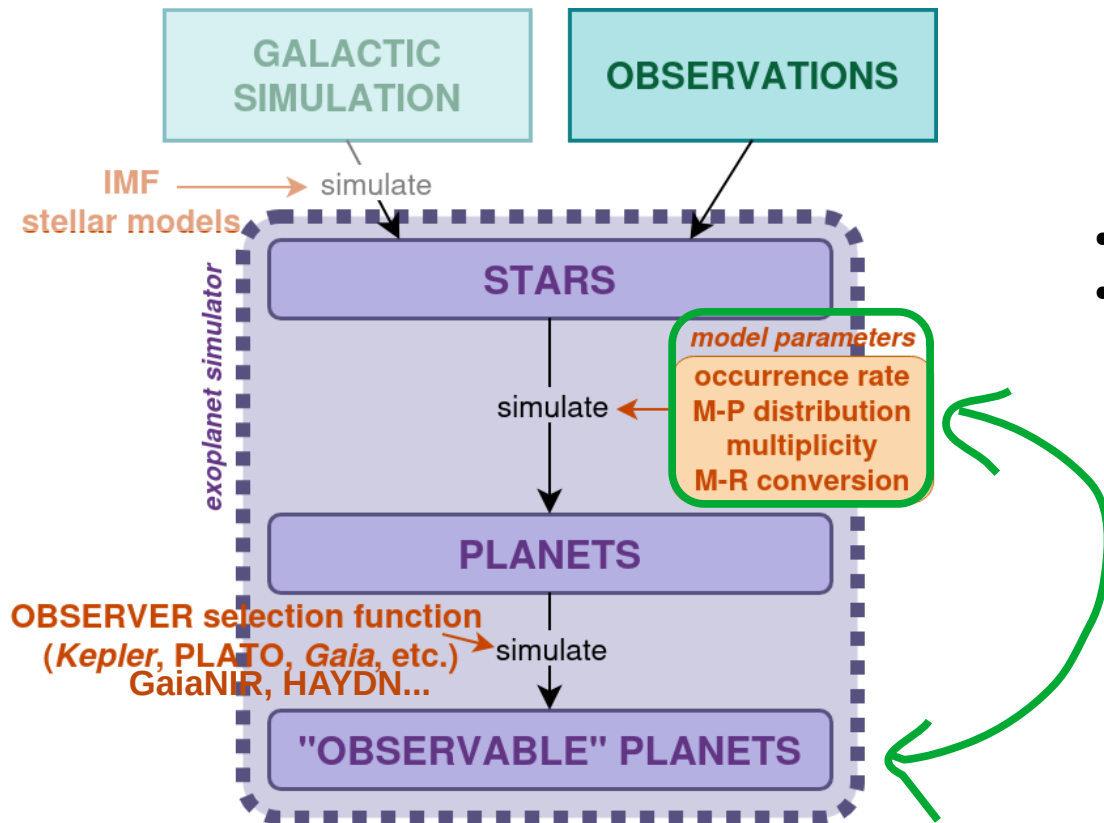


Comparing with observations: adjust model parameters

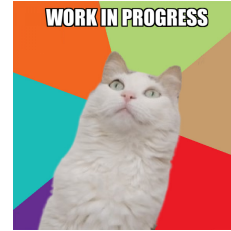


Simulation-Based Inference (SBI)

- Variation of the model parameters
- Learn input parameters-output relation

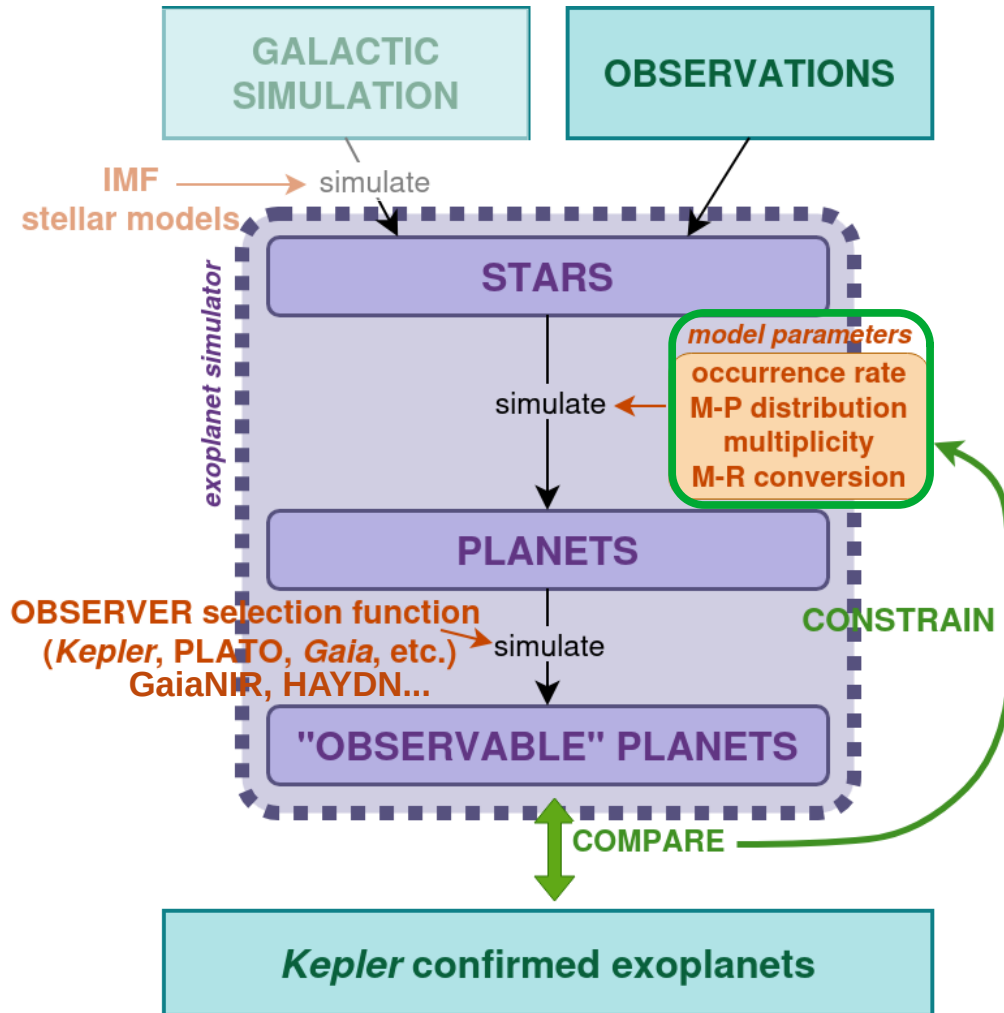


Comparing with observations: adjust model parameters

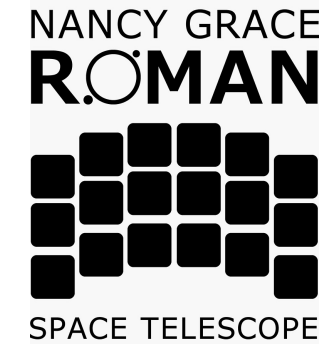
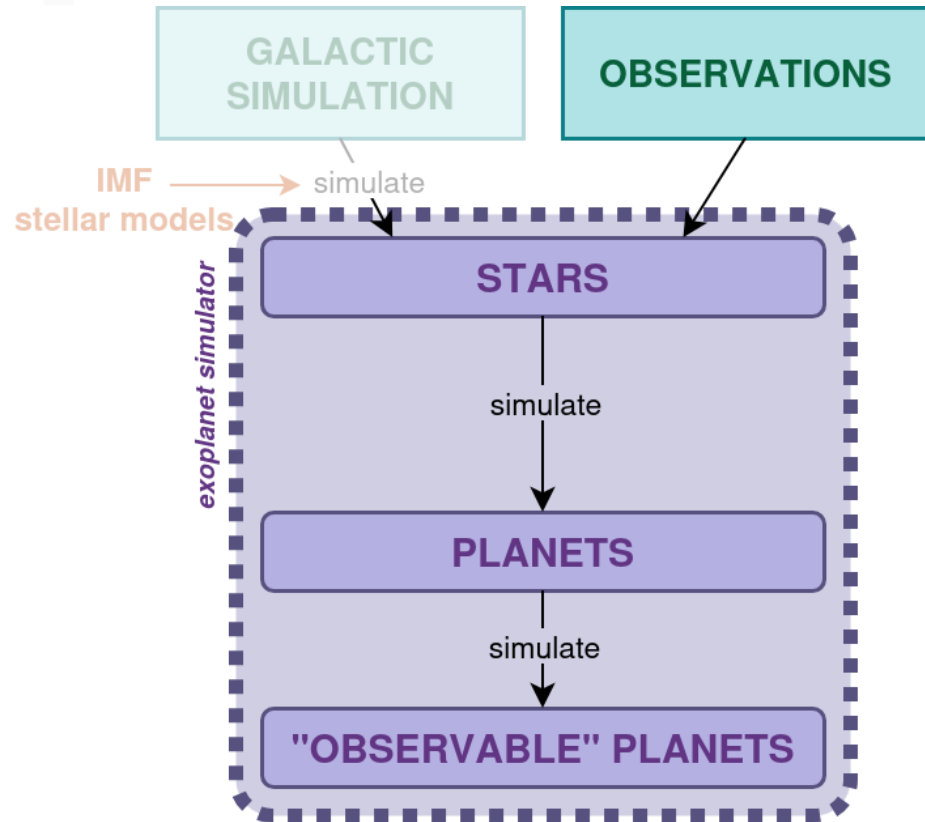


Simulation-Based Inference (SBI)

- Variation of the model parameters
- Learn input parameters-output relation
- Infer model parameters giving the closest output population

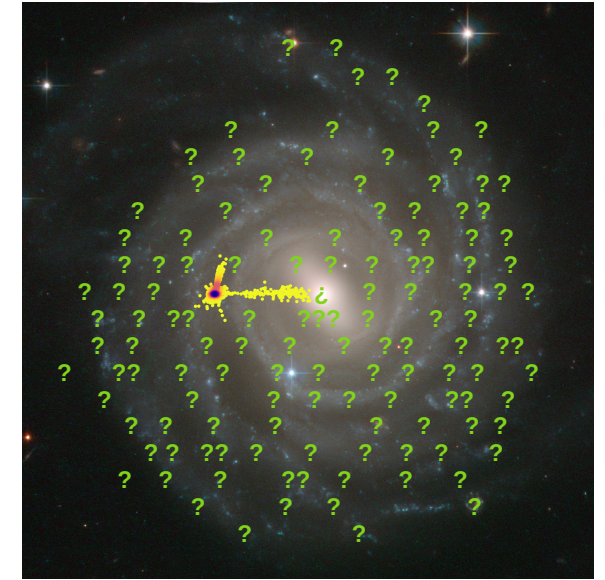


Next step: predict yields of different missions

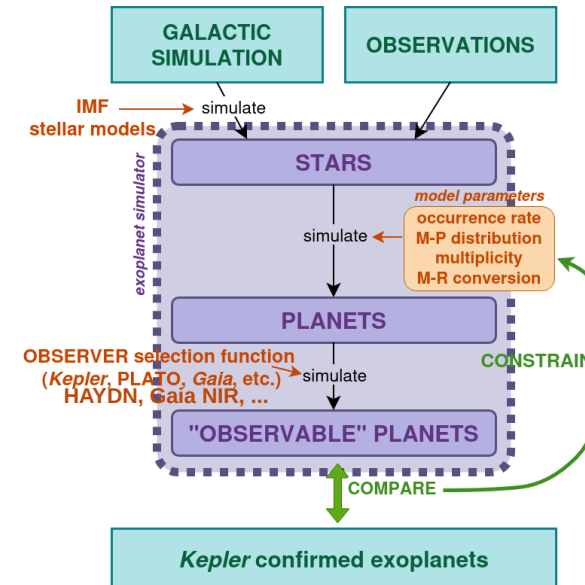


Summary

- Goal: study exoplanet distribution from a Galactic point of view (Padois et al., 2025)
- Created an hybrid exoplanet population synthesis framework
- Can be applied to simulated as well as observed stellar populations



- Apply our framework on Kepler field → **improvements**
- Future: Predictions for future missions → **refine planetary formation models parameters**



Modelling the Milky Way exoplanet population

Chloé Padois, Friedrich Anders, Daniel del Ser, et al.



MWGaiaDN

This project is a Horizon Europe Marie Skłodowska-Curie Actions Doctoral Network funded under grant agreement no. 101072454.

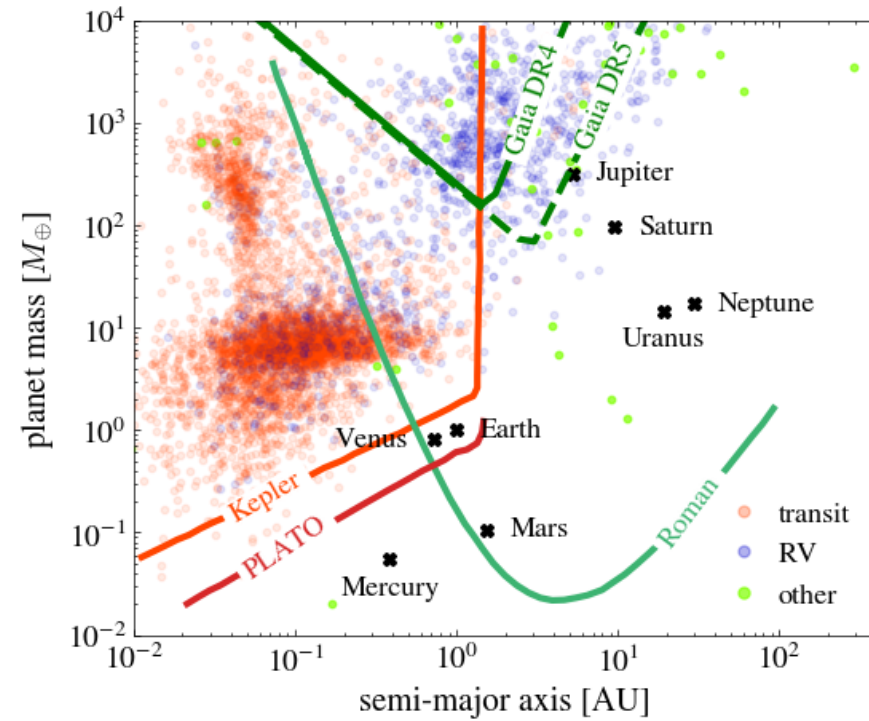


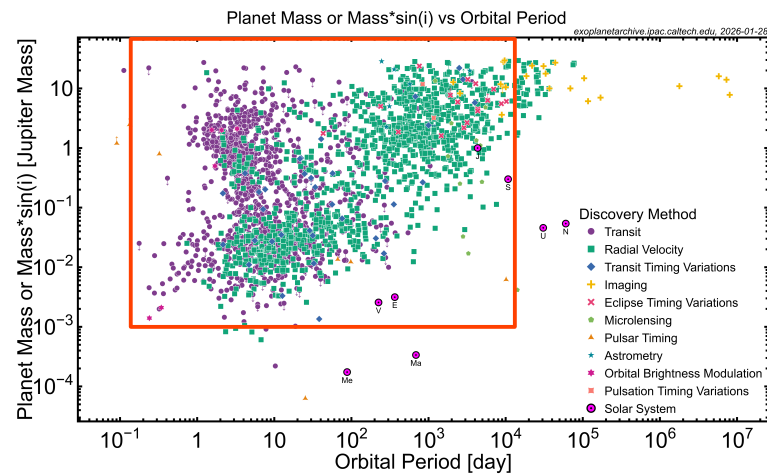
UNIVERSITAT DE
BARCELONA



Merci!

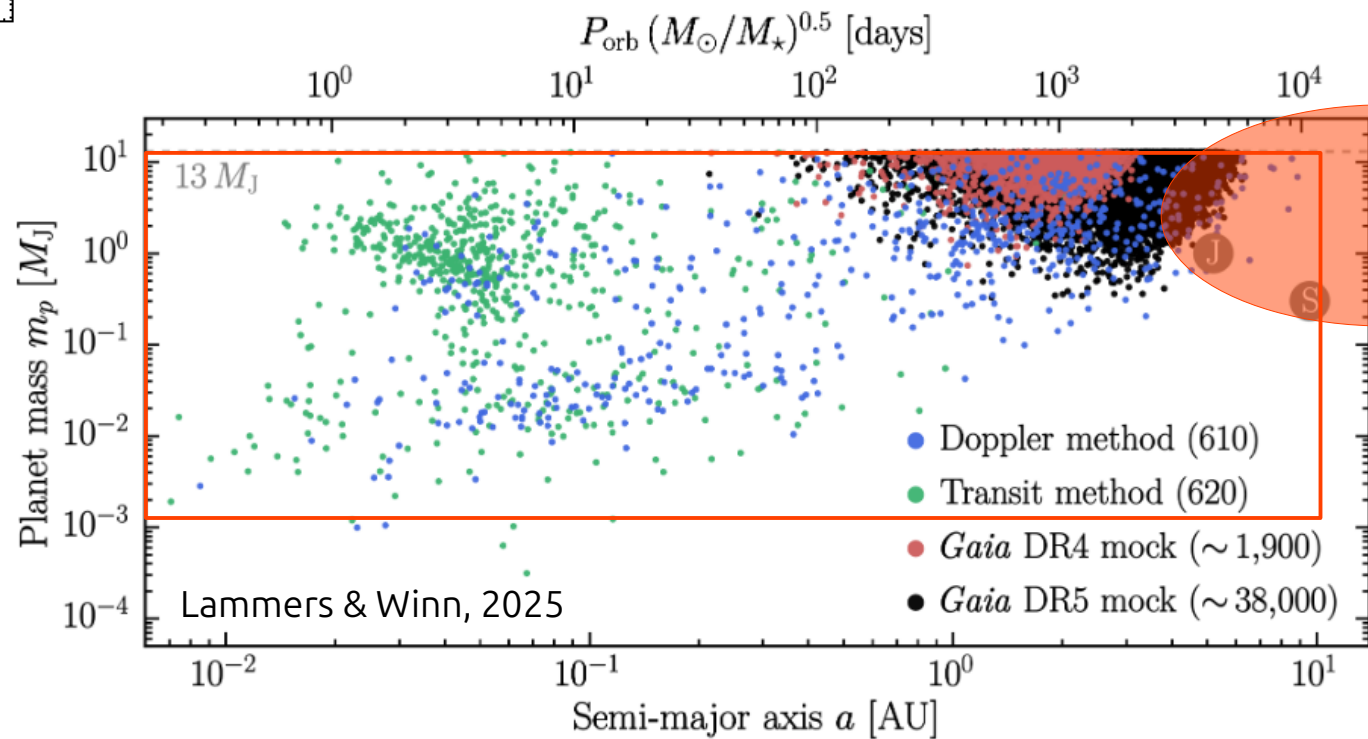
Modelling the Milky Way exoplanet population



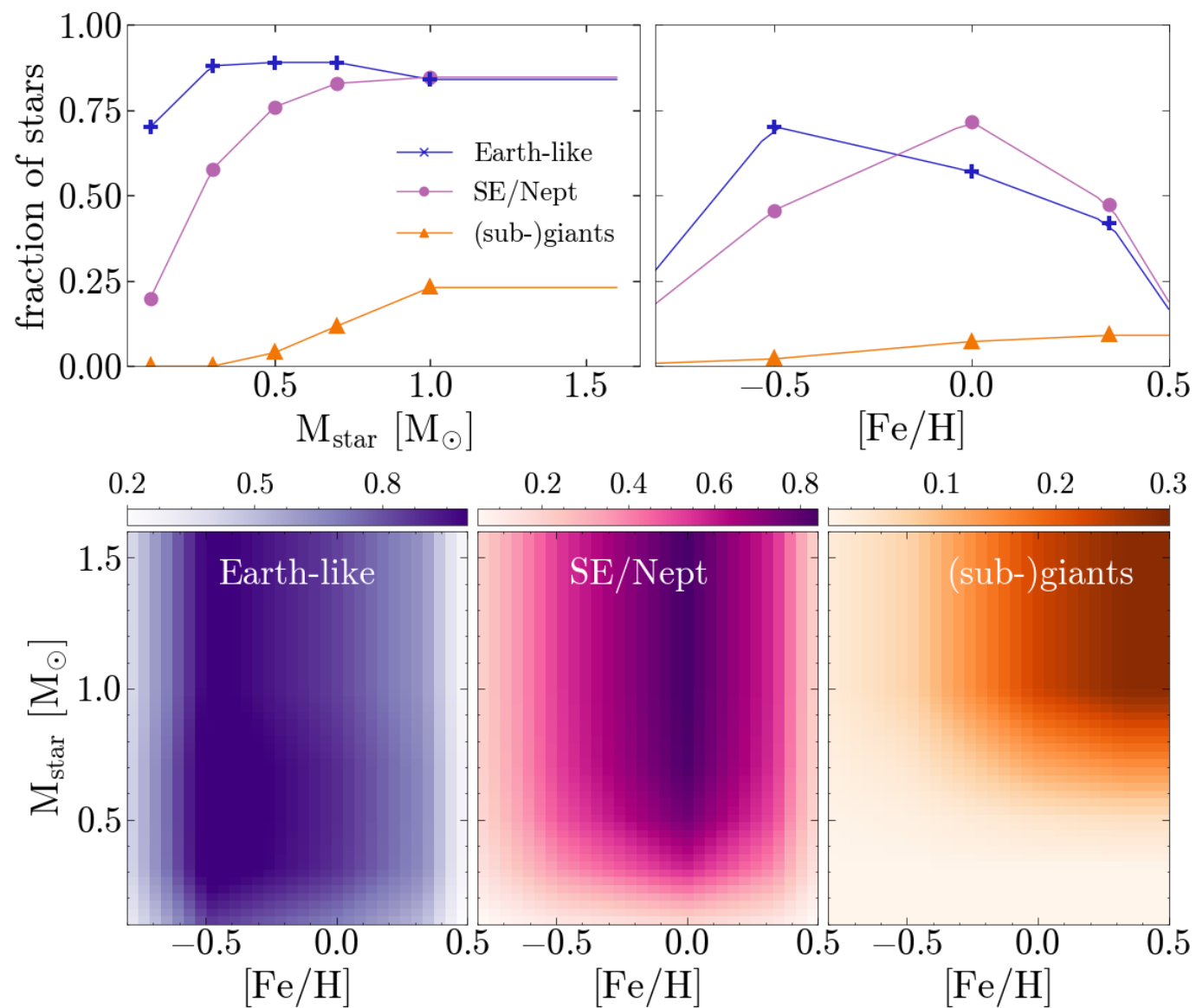


On the Exoplanet Yield of *Gaia* Astrometry

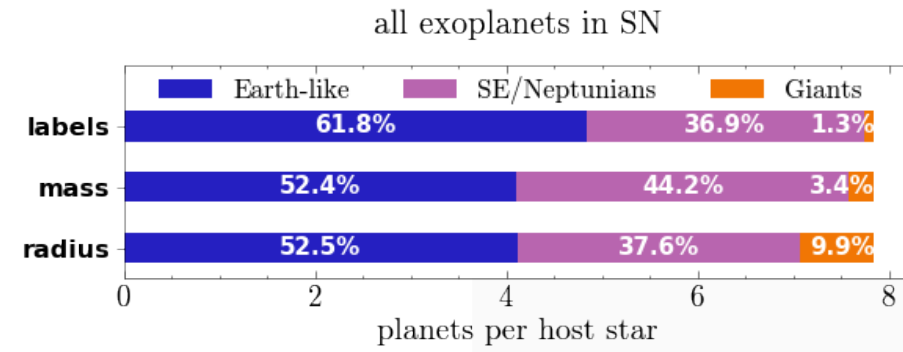
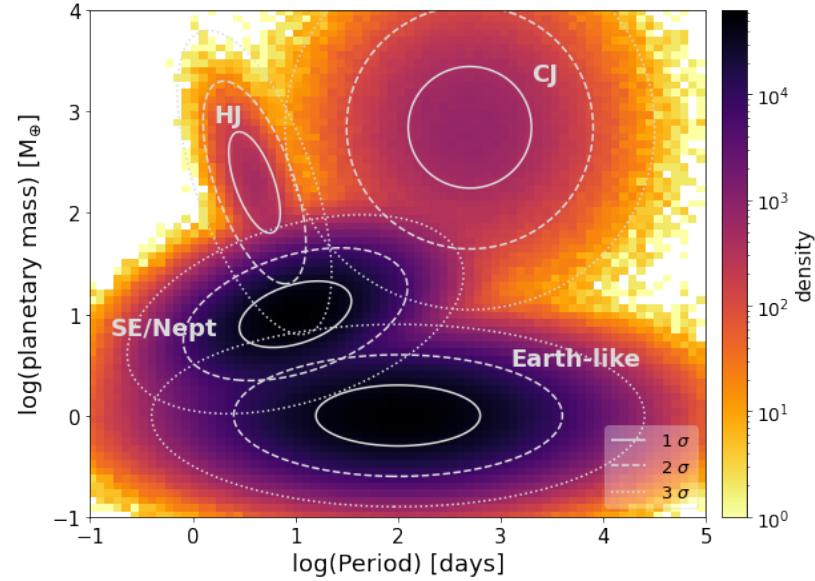
CALEB LAMMERS ¹ AND JOSHUA N. WINN ¹



Extra slides:

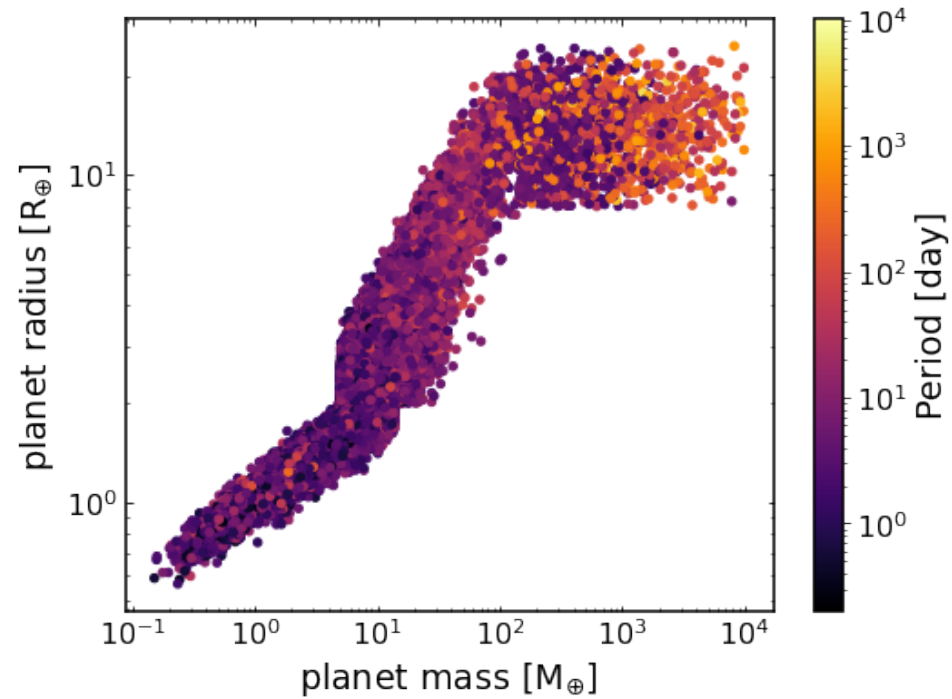


Extra slides:



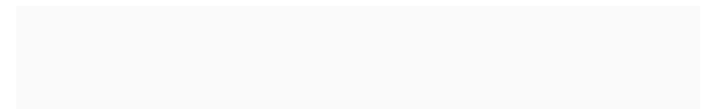
	Earth-like	SE/nept	Giants
radius	$R < 1.25 R_{\text{earth}}$	$1.25 < R < 4$	$4 < R$
mass	$M < 2 M_{\text{earth}}$	$2 < M < 30$	$30 < M$

Extra slides:



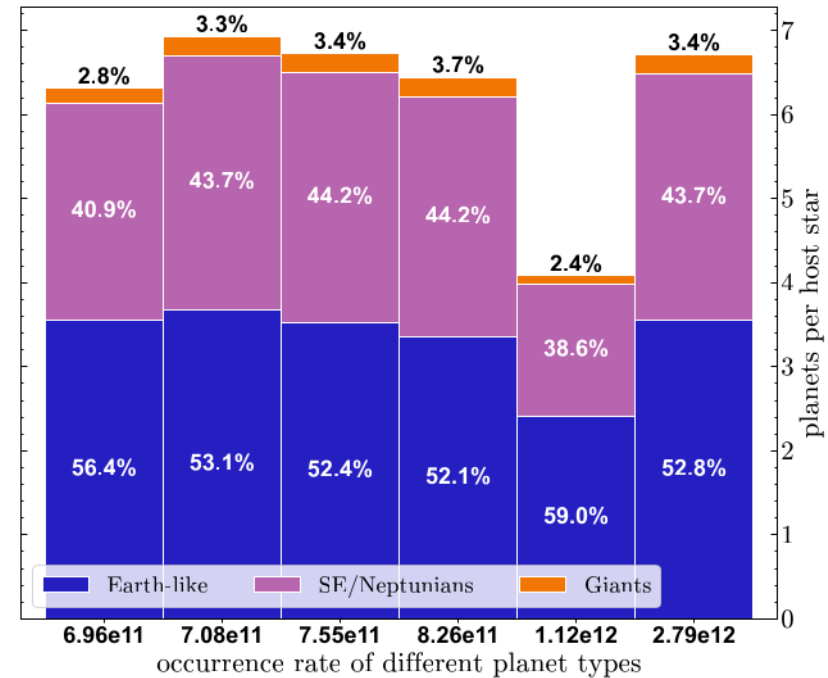
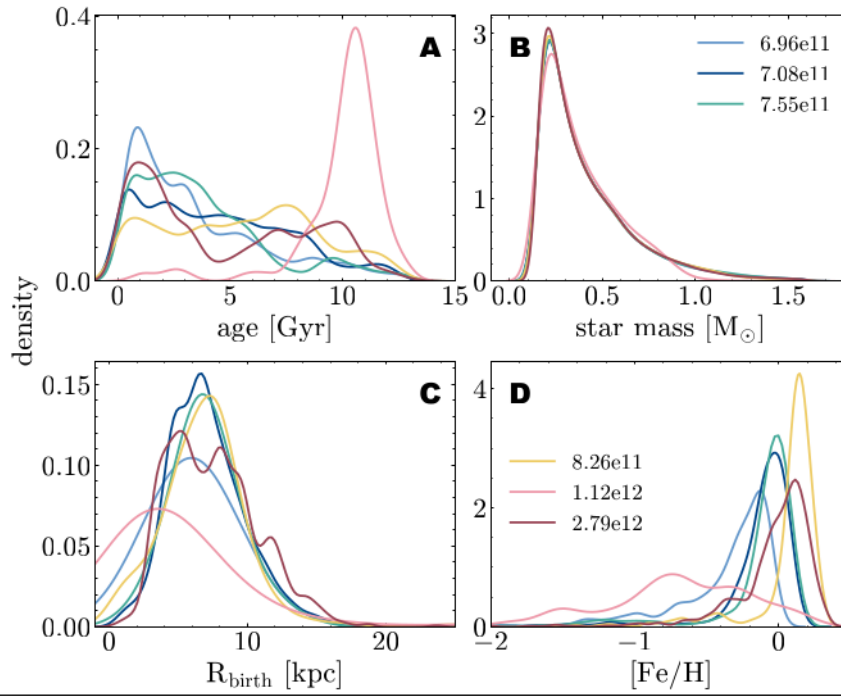
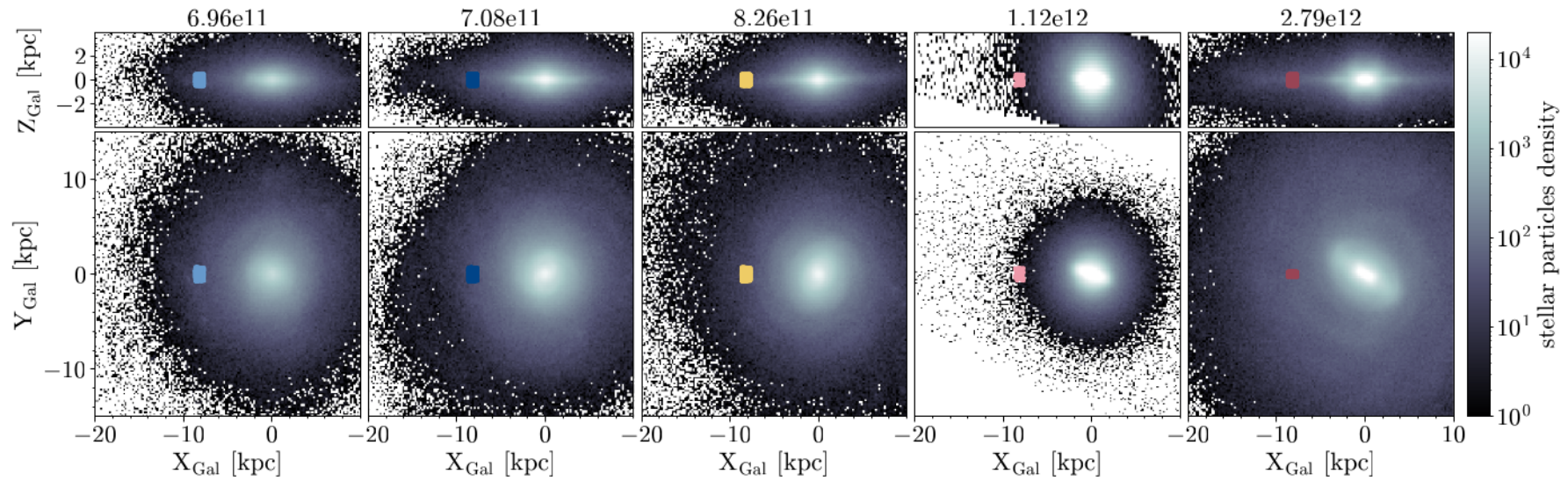
Slopes from Parc+24

→ for *Kepler* "detectable" planets



Modelling the Milky Way exoplanet population

Extra slides:



Modelling the Milky Way exoplanet population

Extra slides:

